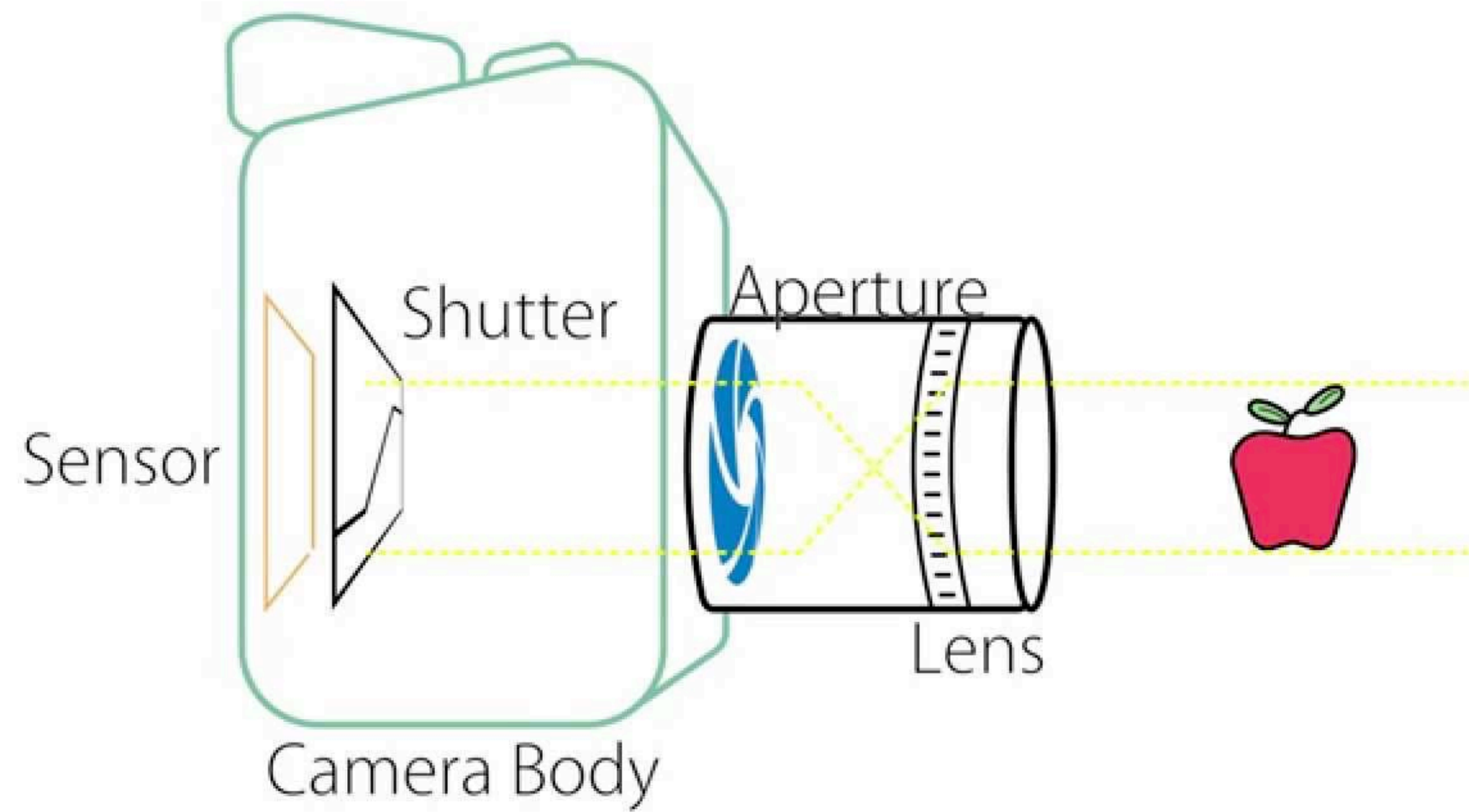
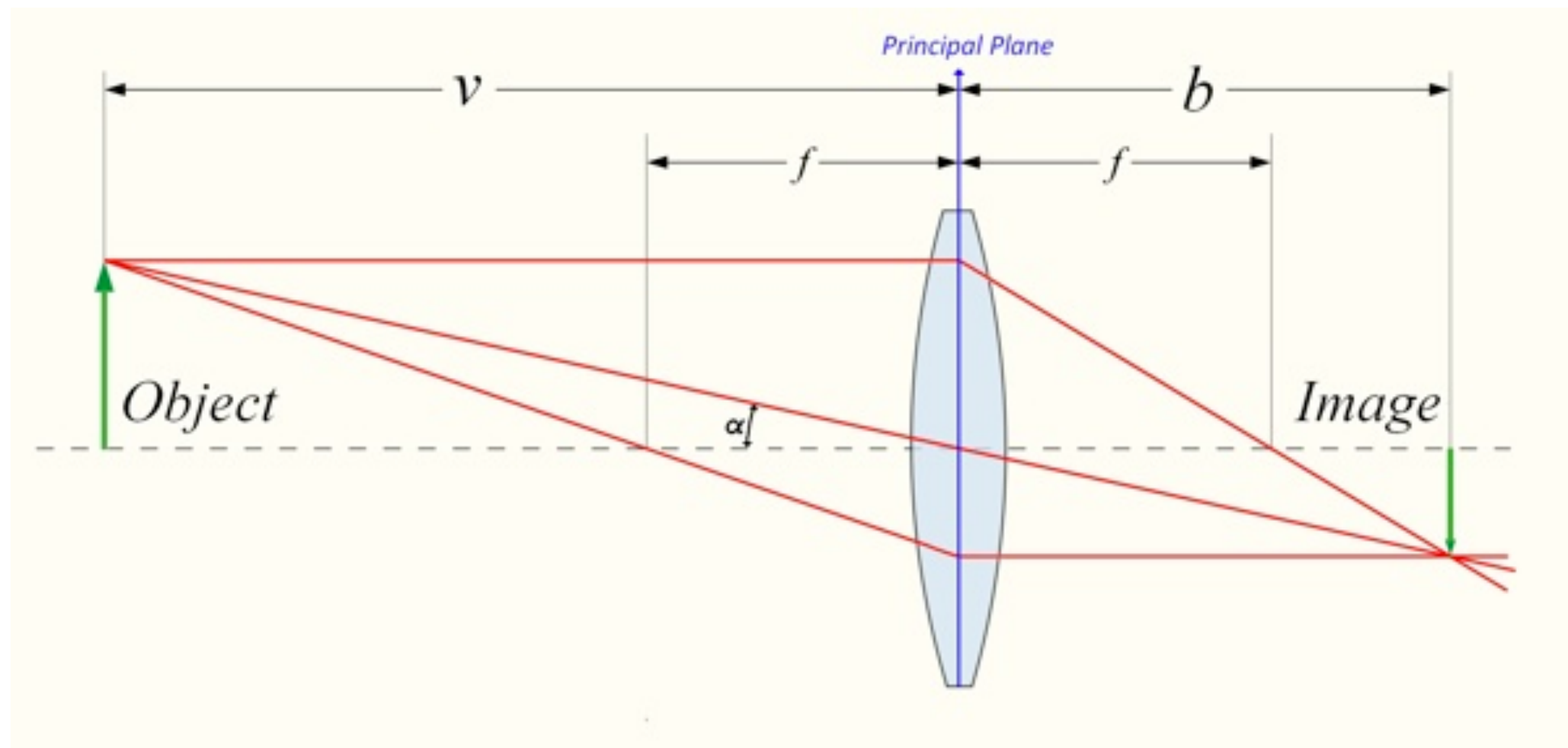
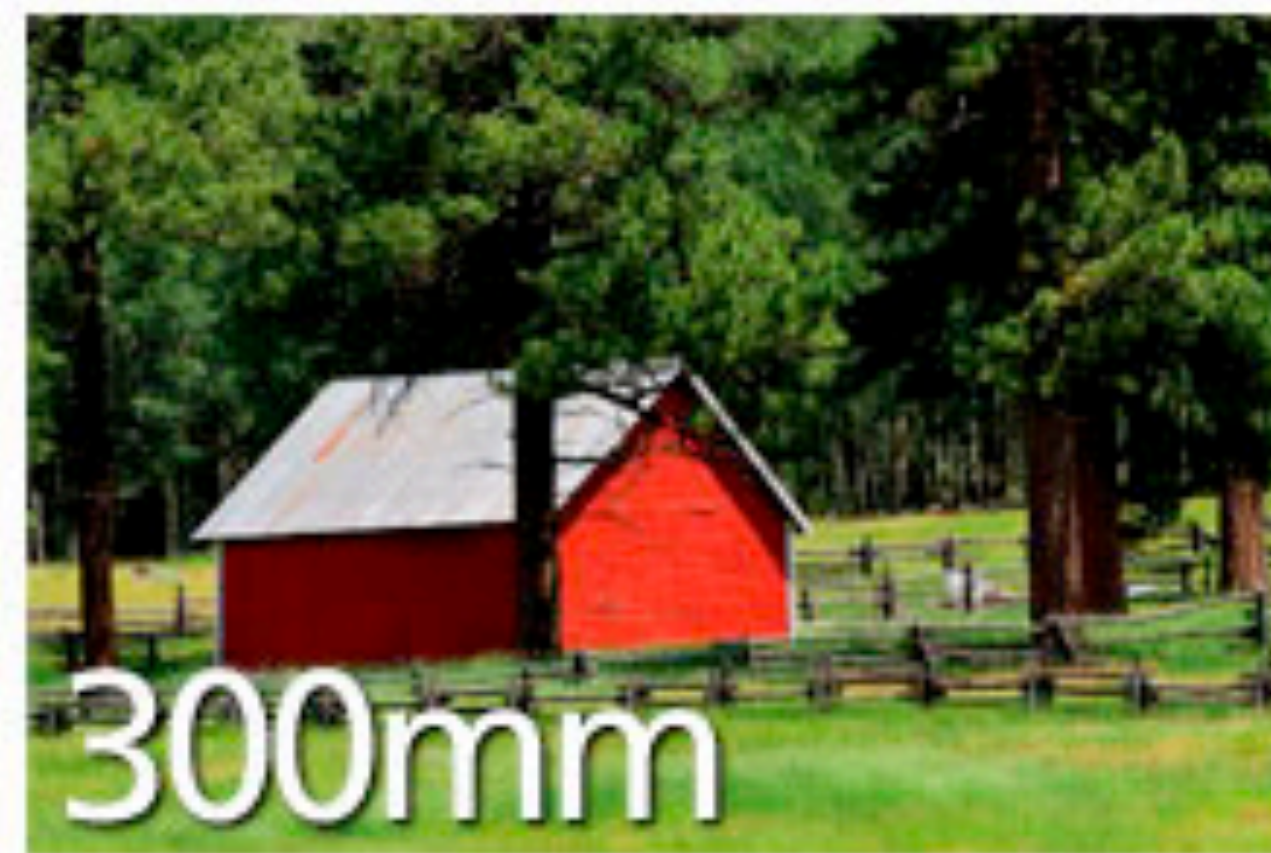
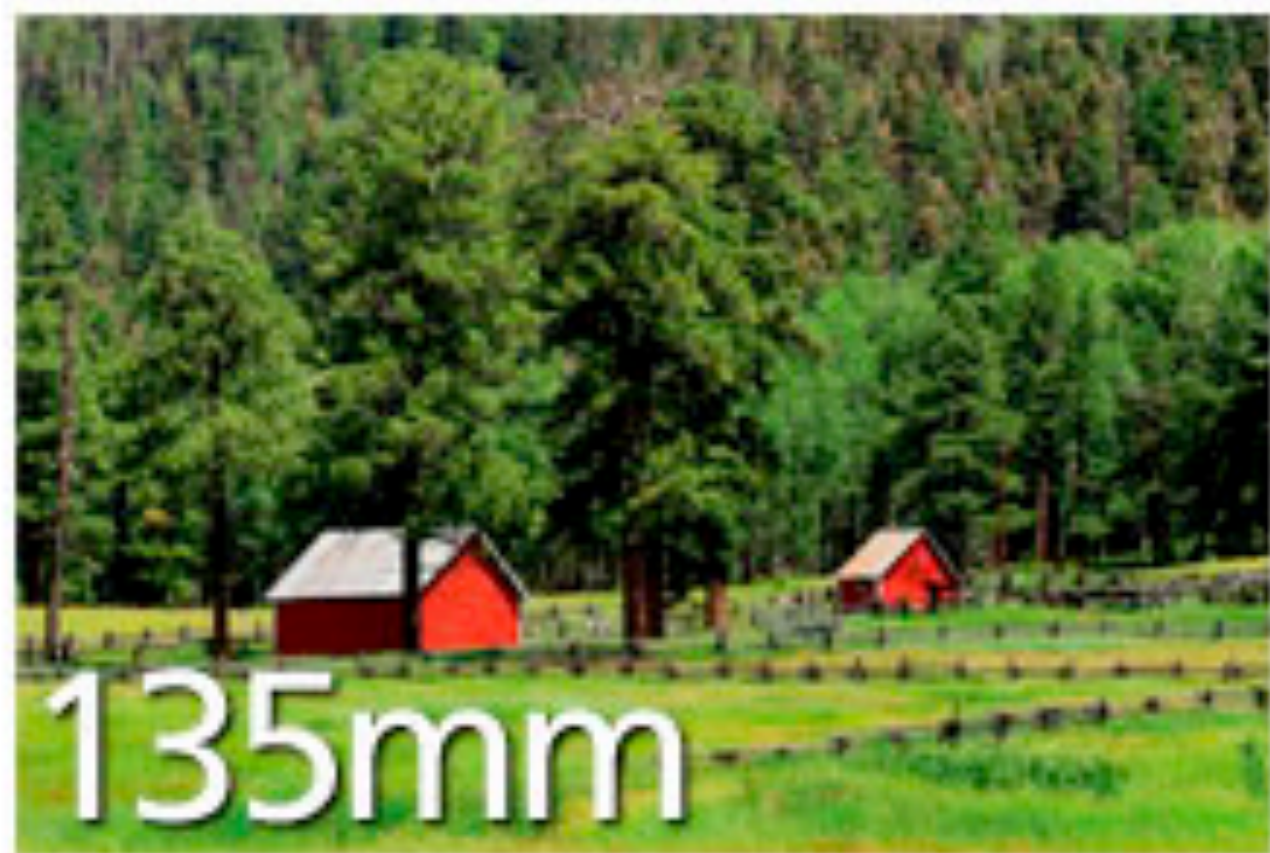
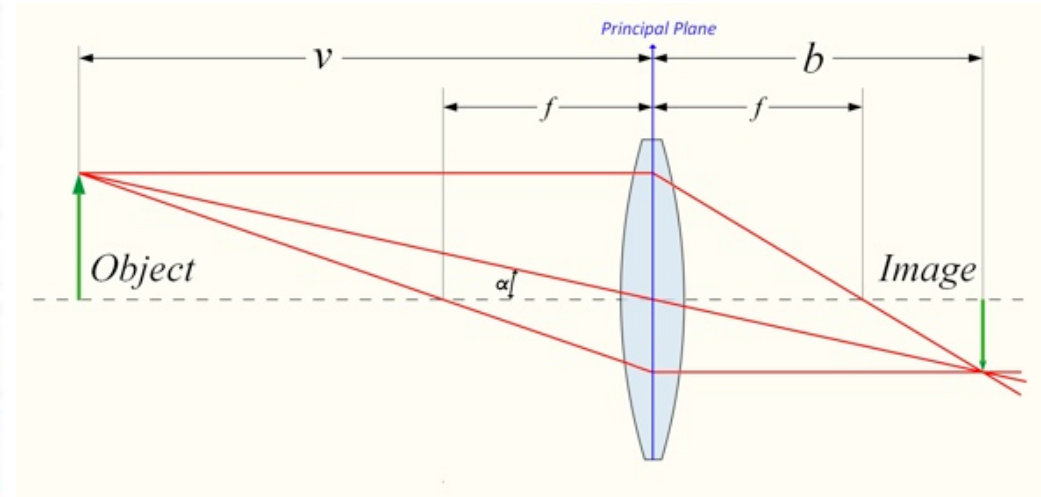


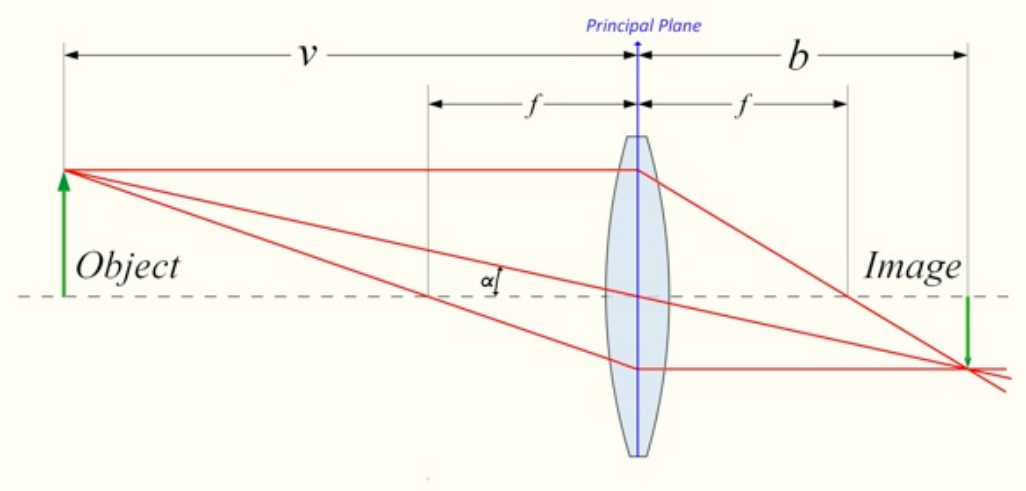
# **Robotica autonoma en entornos hostiles**

**VISION 3D**

# Imágenes Digitales







DX-FORMAT is equivalent to a 1.5x crop compared to FX format

10.5MM Full Frame Fisheye

16MM

35MM

135MM

FIELD OF VIEW

12°

46°

84°

180°



200mm



300mm



16mm

the  
struggle

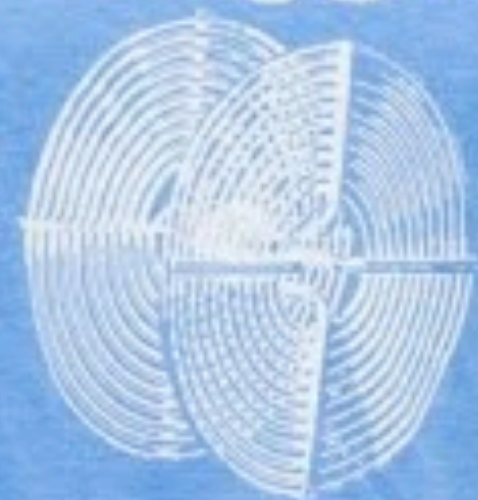


is reel



50mm

the  
struggle

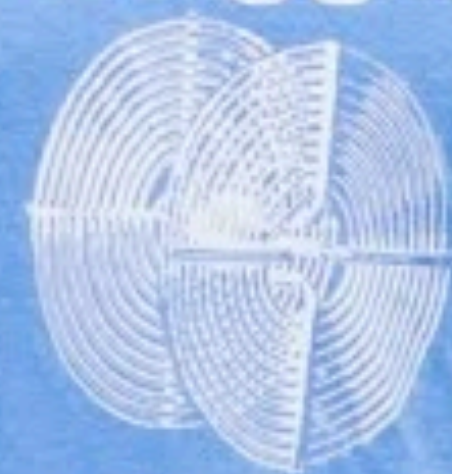


is reel



300mm

the  
struggle



is reel

## SHUTTER SPEED



## MOTION

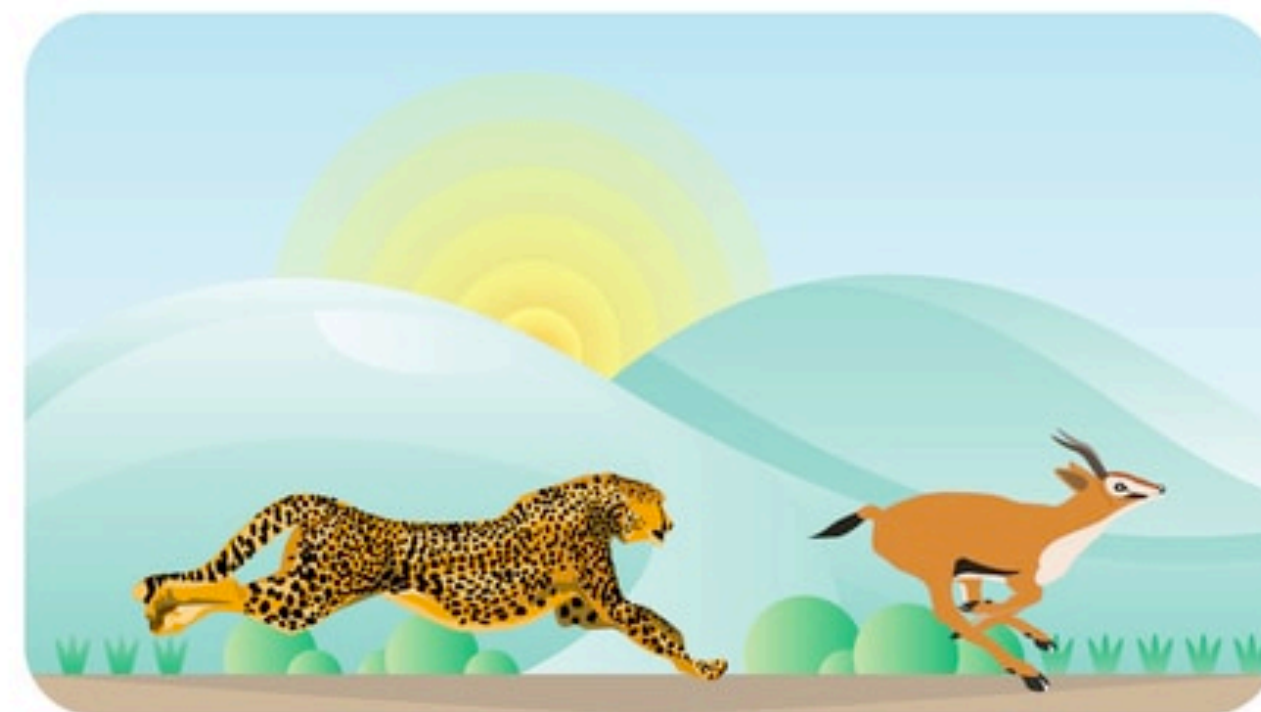
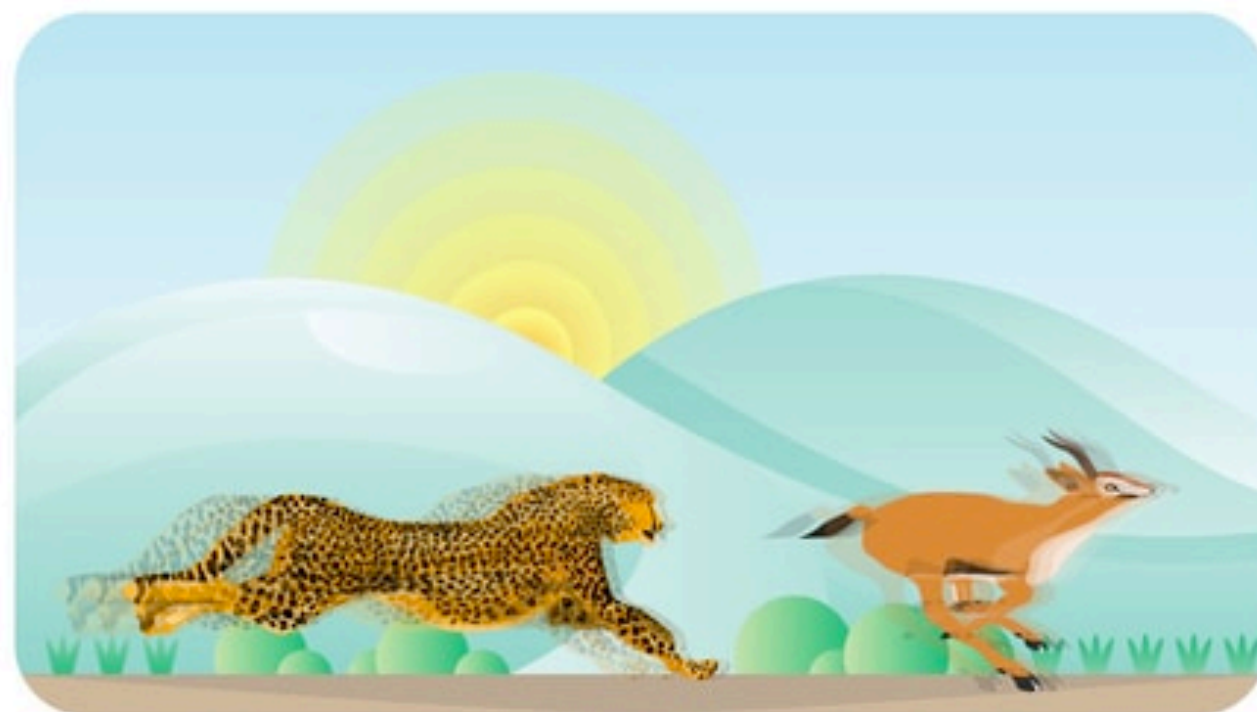
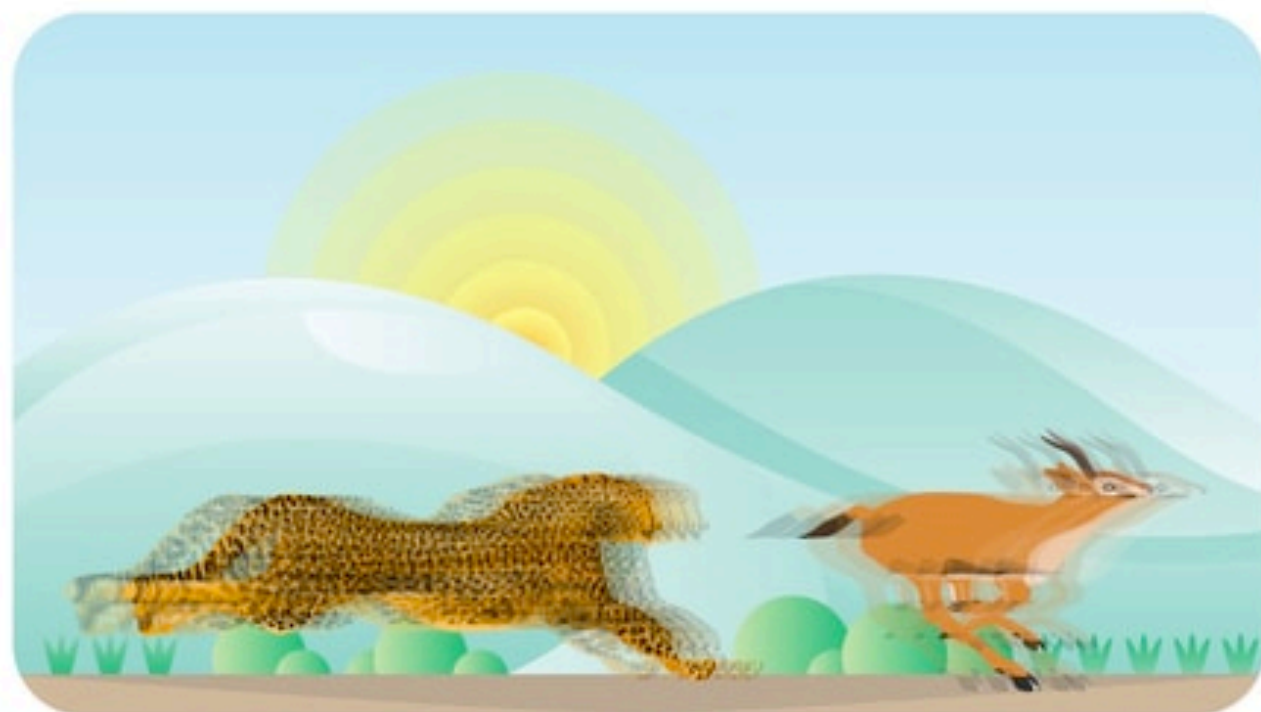
# Shutter Speed and Motion

SLOWER

FASTER

Blurry Motion

Freeze Motion



Shutter Speed: 1/1000



Shutter Speed: 1/125



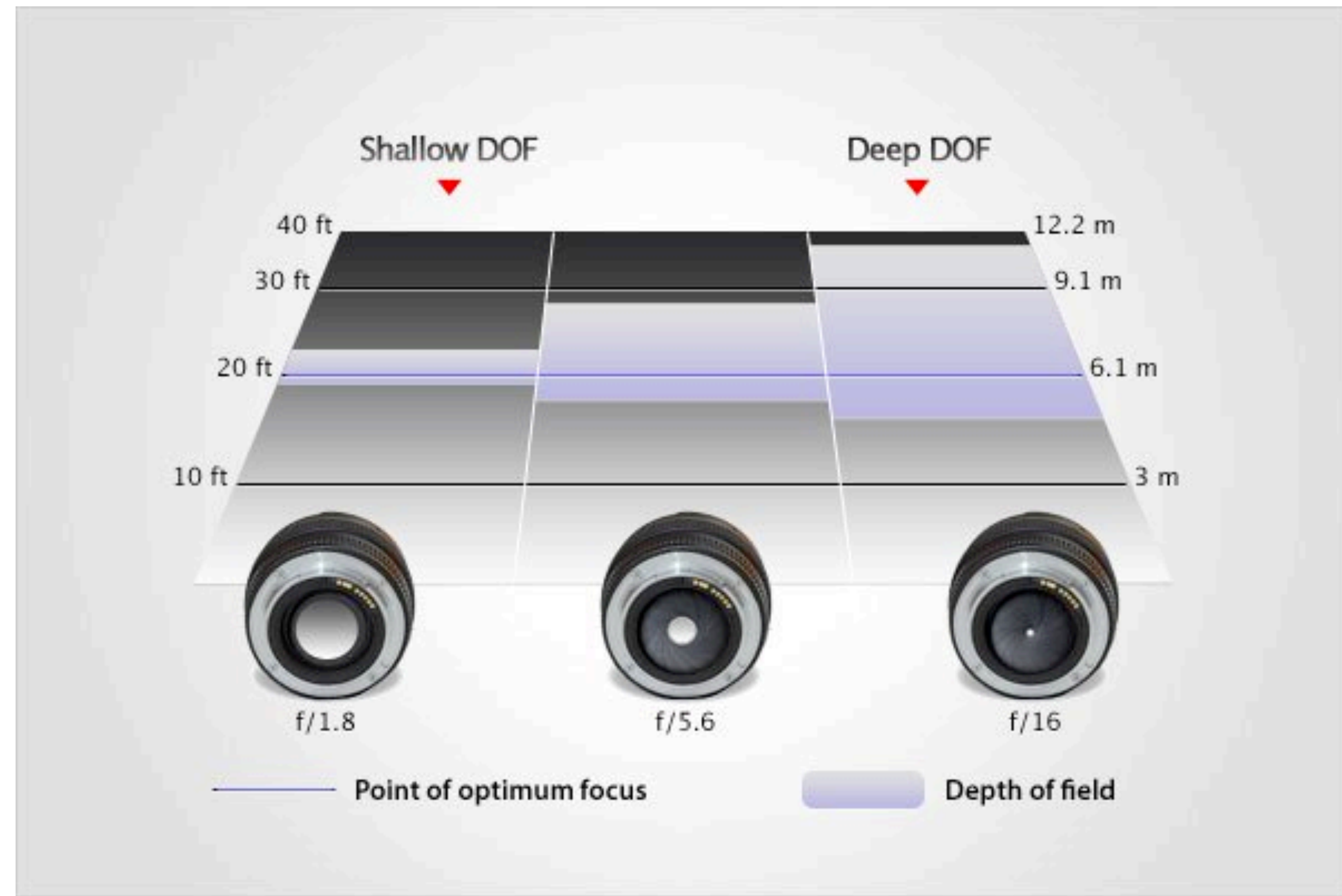
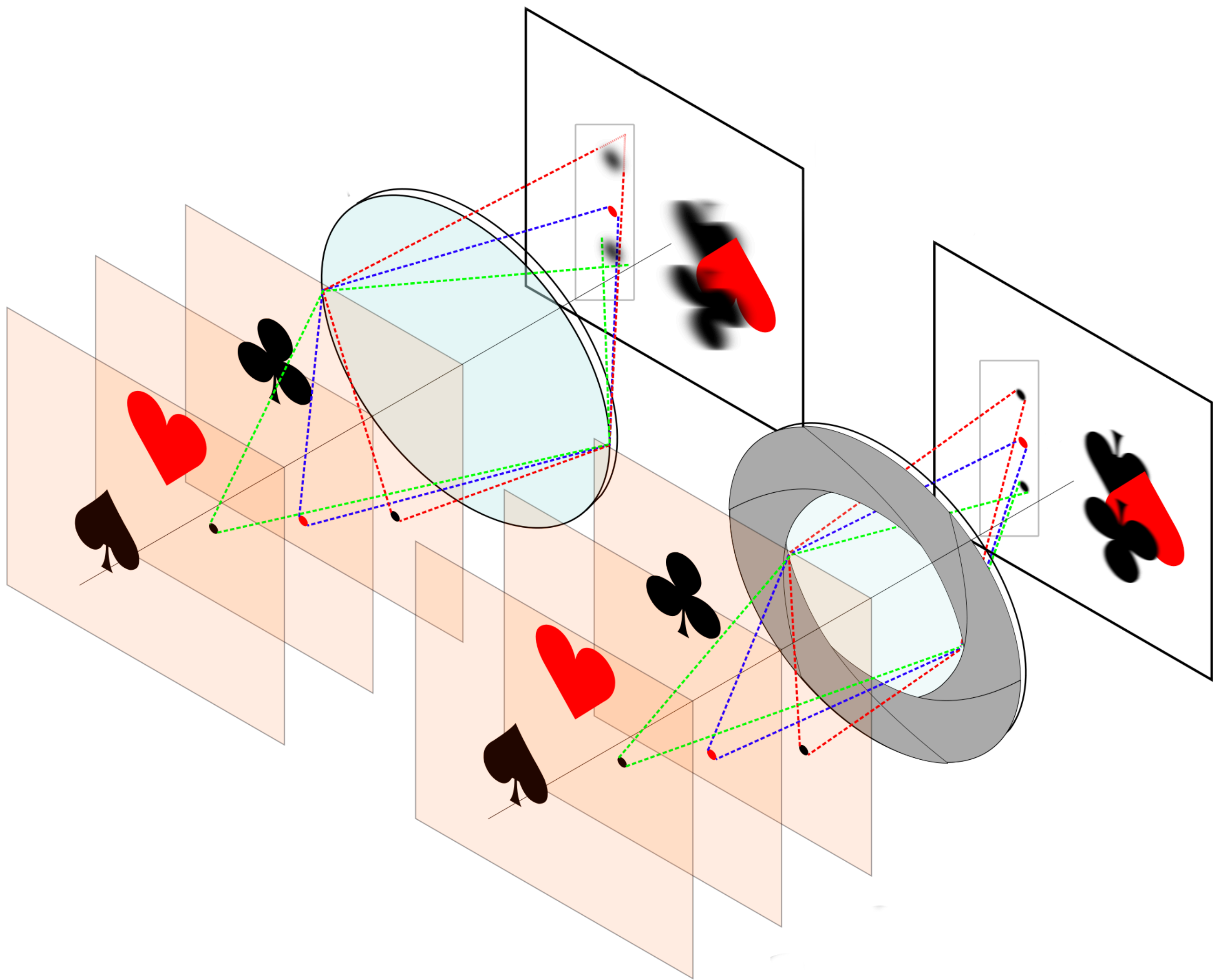
Shutter Speed: 1/25



Shutter Speed: 1/15









f/1.4



f/2



f/2.8



f/4



f/5.6



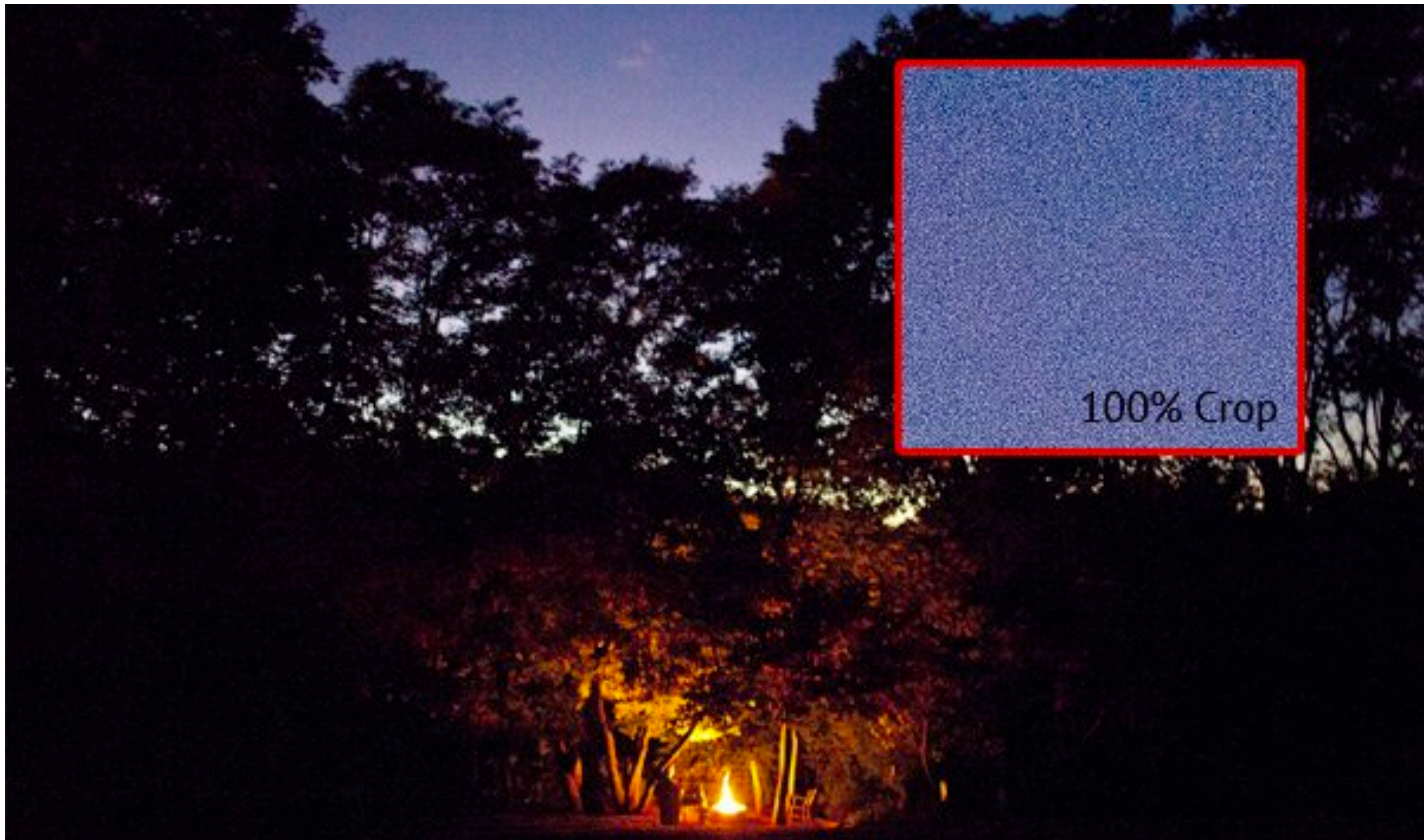
f/8



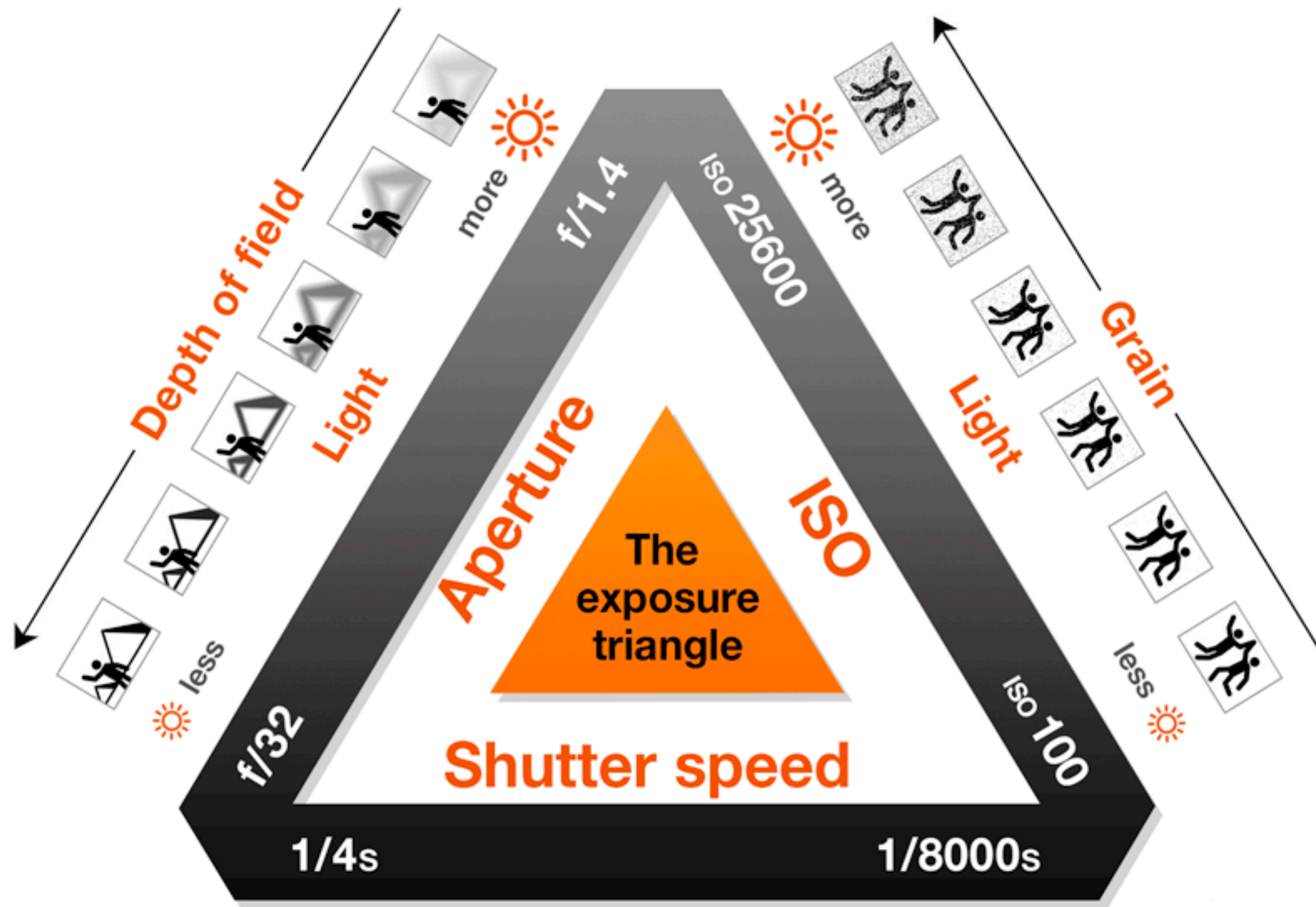
f/11



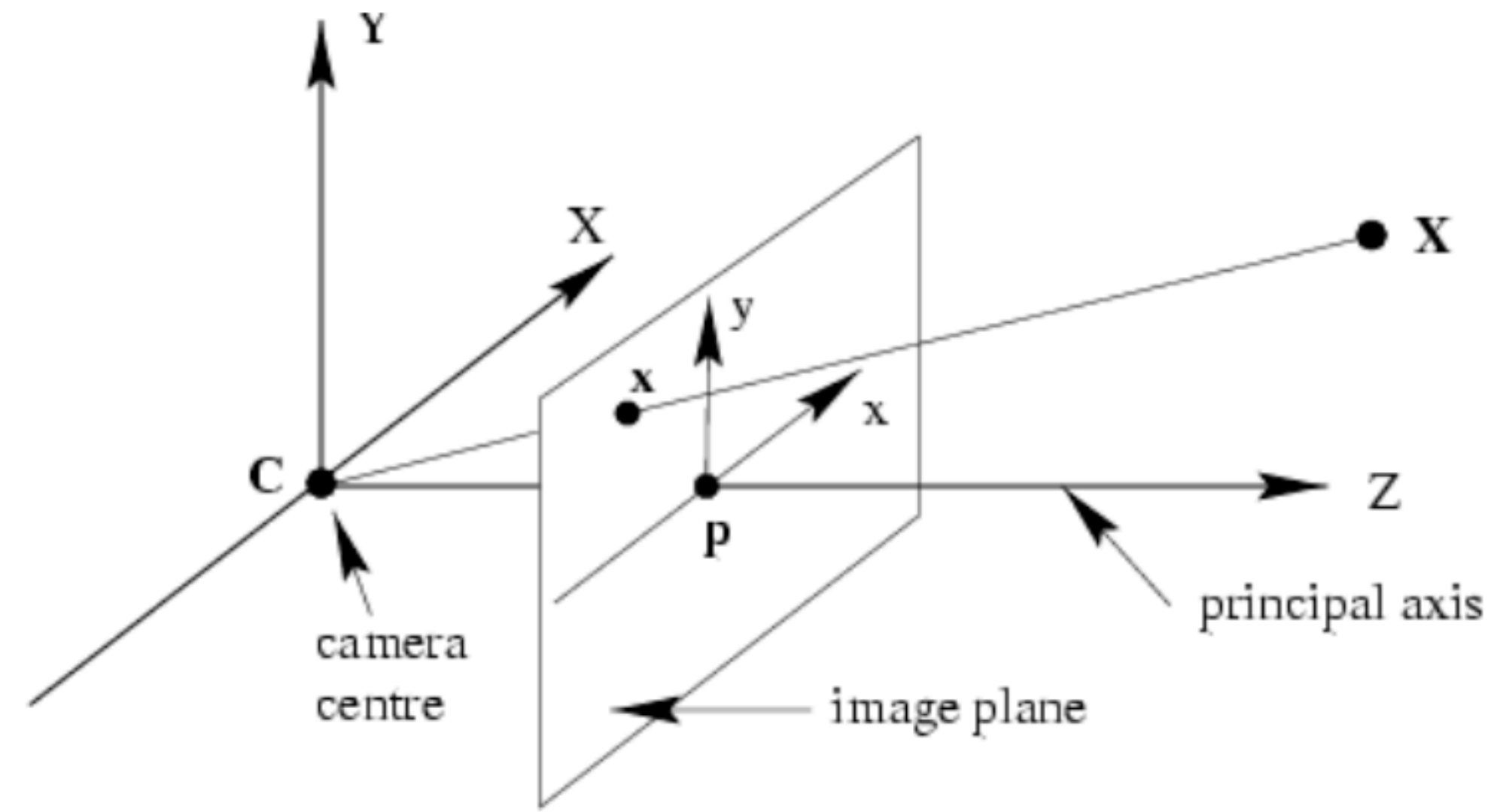
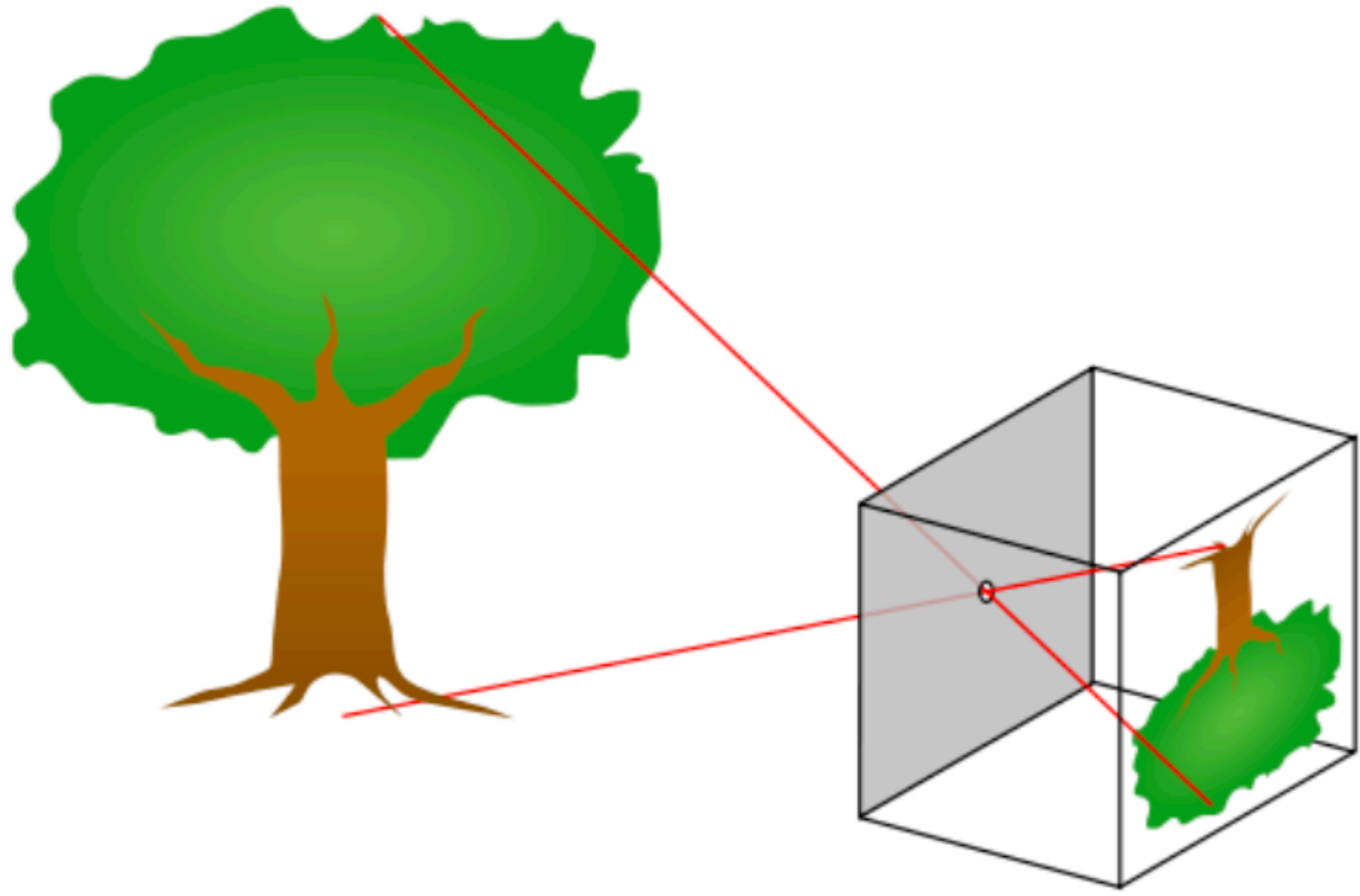
f/16



100% Crop



**Como modelamos una cámara?**



Model

Projection (Source: Wikipedia)

The “pinhole” camera (French: sténopé):

- ▶ Ideal model with an aperture reduced to a single point.
- ▶ No account for blur of out of focus objects, nor for the lens geometric distortion.

## Central projection in camera coordinate frame

- ▶ Rays from  $C$  are the same:  $\vec{C}_x = \lambda \vec{C}\vec{X}$
- ▶ In the camera coordinate frame  $CXYZ$ :

$$\begin{pmatrix} x \\ y \\ f \end{pmatrix} = \lambda \begin{pmatrix} X \\ Y \\ Z \end{pmatrix}$$

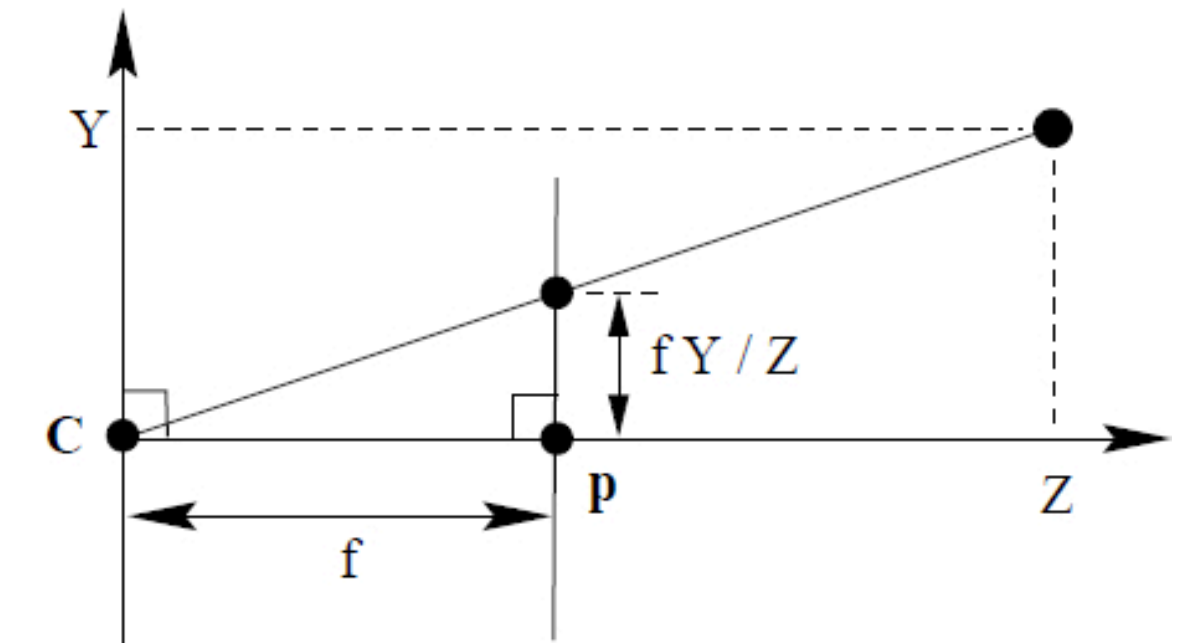
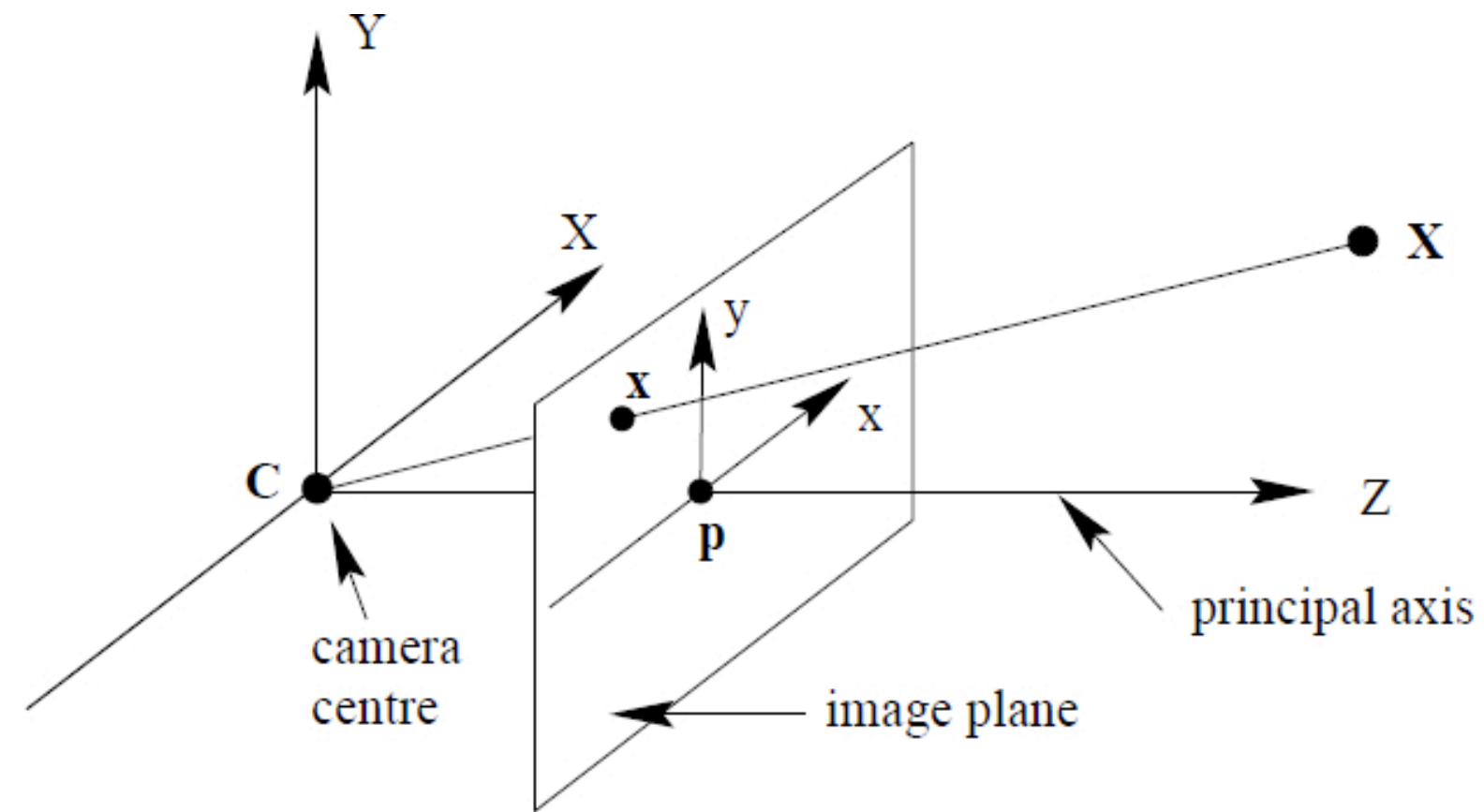
- ▶ Thus  $\lambda = f/Z$  and

$$\begin{pmatrix} x \\ y \end{pmatrix} = f \begin{pmatrix} X/Z \\ Y/Z \end{pmatrix}$$

- ▶ In pixel coordinates:

$$\begin{pmatrix} u \\ v \end{pmatrix} = \begin{pmatrix} \alpha x + c_x \\ \alpha y + c_y \end{pmatrix} = \begin{pmatrix} (\alpha f)X/Z + c_x \\ (\alpha f)Y/Z + c_y \end{pmatrix}$$

- ▶  $\alpha f$ : focal length *in pixels*,  $(c_x, c_y)$ : position of principal point  $P$  in pixels.



## Calibration matrix

- ▶ Let us get back to the projection equation:

$$\begin{pmatrix} u \\ v \end{pmatrix} = \begin{pmatrix} fX/Z + c_x \\ fY/Z + c_y \end{pmatrix} = \frac{1}{Z} \begin{pmatrix} fX + c_x Z \\ fY + c_y Z \end{pmatrix}$$

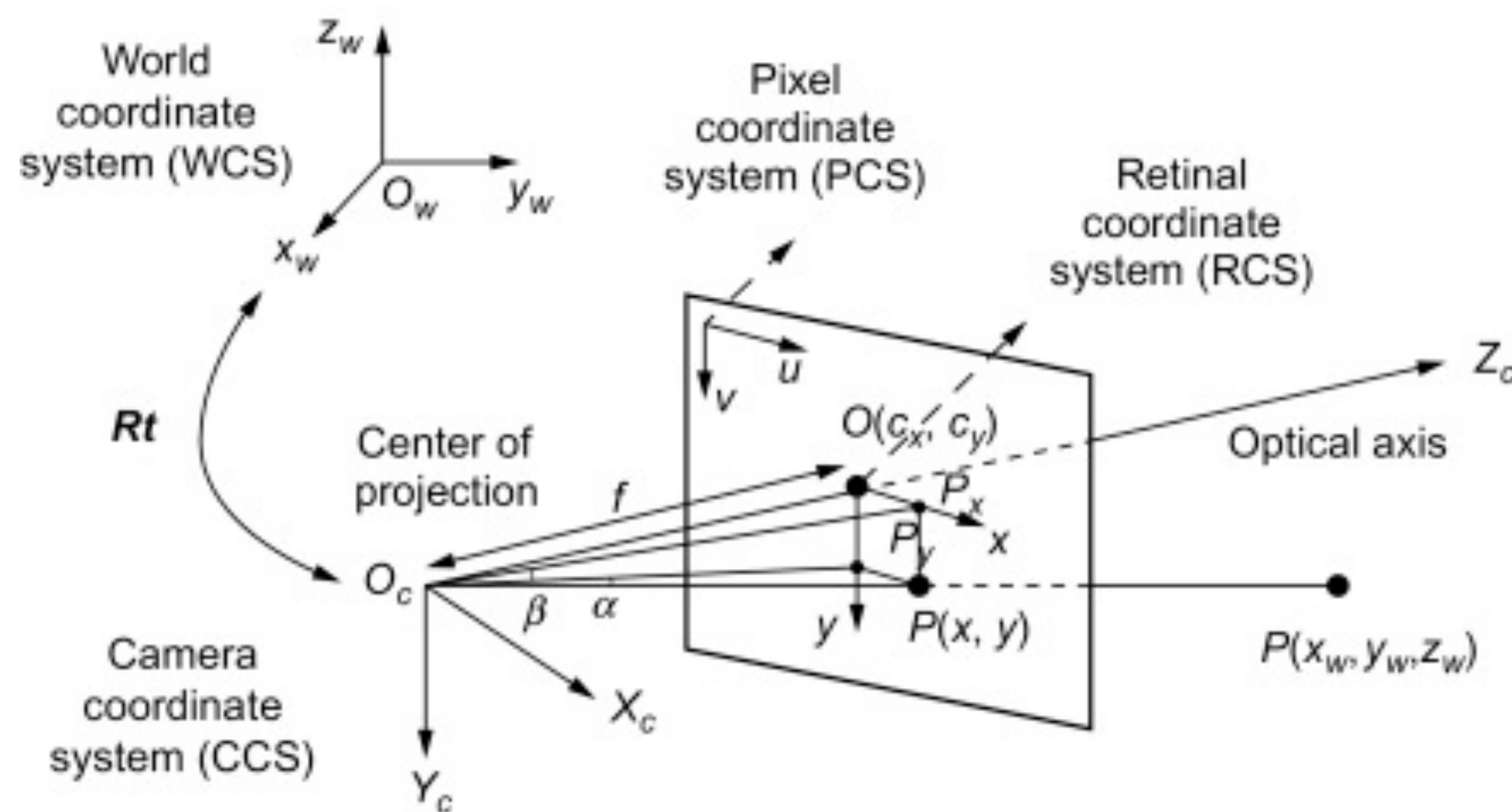
(replacing  $\alpha f$  by  $f$ )

- ▶ We rewrite:

$$Z \begin{pmatrix} u \\ v \\ 1 \end{pmatrix} := x = \begin{pmatrix} f & c_x \\ & f & c_y \\ & & 1 \end{pmatrix} \begin{pmatrix} X \\ Y \\ Z \end{pmatrix}$$

- ▶ The 3D point being expressed in another orthonormal coordinate frame:

$$x = \begin{pmatrix} f & c_x \\ & f & c_y \\ & & 1 \end{pmatrix} (R \ T) \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix}$$





## Calibration matrix

- ▶ Let us get back to the projection equation:

$$\begin{pmatrix} u \\ v \end{pmatrix} = \begin{pmatrix} fX/Z + c_x \\ fY/Z + c_y \end{pmatrix} = \frac{1}{Z} \begin{pmatrix} fX + c_x Z \\ fY + c_y Z \end{pmatrix}$$

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- ▶ The 3D point being expressed in another orthonormal coordinate frame:

$$x = \begin{pmatrix} f & c_x \\ & f & c_y \\ & & 1 \end{pmatrix} (R \ T) \begin{pmatrix} X \\ Y \\ Z \\ 1 \end{pmatrix}$$

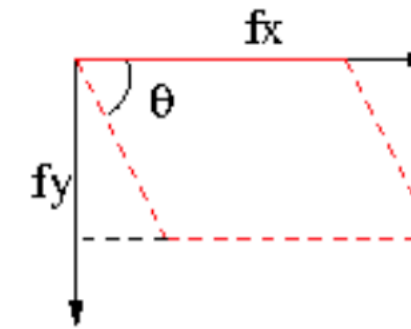
- ▶ The (internal) **calibration matrix** ( $3 \times 3$ ) is:

$$K = \begin{pmatrix} f & c_x \\ & f & c_y \\ & & 1 \end{pmatrix}$$

- ▶ The **projection matrix** ( $3 \times 4$ ) is:

$$P = K (R \ T)$$

- ▶ If pixels are trapezoids, we can generalize  $K$ :



$$K = \begin{pmatrix} f_x & s & c_x \\ & f_y & c_y \\ & & 1 \end{pmatrix} \text{ (with } s = -f_x \cotan \theta \text{)}$$

## Function Documentation

### ◆ calibrateCamera() [1/2]

```
double cv::calibrateCamera ( InputArrayOfArrays objectPoints,
                             InputArrayOfArrays imagePoints,
                             Size imageSize,
                             InputOutputArray cameraMatrix,
                             InputOutputArray distCoeffs,
                             OutputArrayOfArrays rvecs,
                             OutputArrayOfArrays tvecs,
                             OutputArray stdDeviationsIntrinsics,
                             OutputArray stdDeviationsExtrinsics,
                             OutputArray perViewErrors,
                             int flags = 0 ,
                             TermCriteria criteria = TermCriteria(TermCriteria::COUNT+TermCriteria::EPS, 30, DBL_EPSILON)
)
```

#### Python:

```
cv.calibrateCamera( objectPoints, imagePoints, imageSize, cameraMatrix, distCoeffs[, rvecs[, tvecs[, flags[, criteria]]])
```

#### Parameters

##### **objectPoints**

In the new interface it is a vector of vectors of calibration pattern points in the many elements as the number of pattern views. If the same calibration pattern is used in partially occluded patterns or even different patterns in different views. Then plane (thus 0 in the Z-coordinate), if the used calibration pattern is a planar

##### **imagePoints**

In the new interface it is a vector of vectors of the projections of calibration pattern points. `imagePoints[i].size()` and `objectPoints[i].size()` for each `i`, must be equal, respectively

##### **imageSize**

Size of the image used only to initialize the camera intrinsic matrix.

##### **cameraMatrix**

Input/output 3x3 floating-point camera intrinsic matrix  $A = \begin{bmatrix} f_x & 0 & c_x \\ 0 & f_y & c_y \\ 0 & 0 & 1 \end{bmatrix}$ .

**Cómo podemos extraer  
información 3D usando cámaras?**

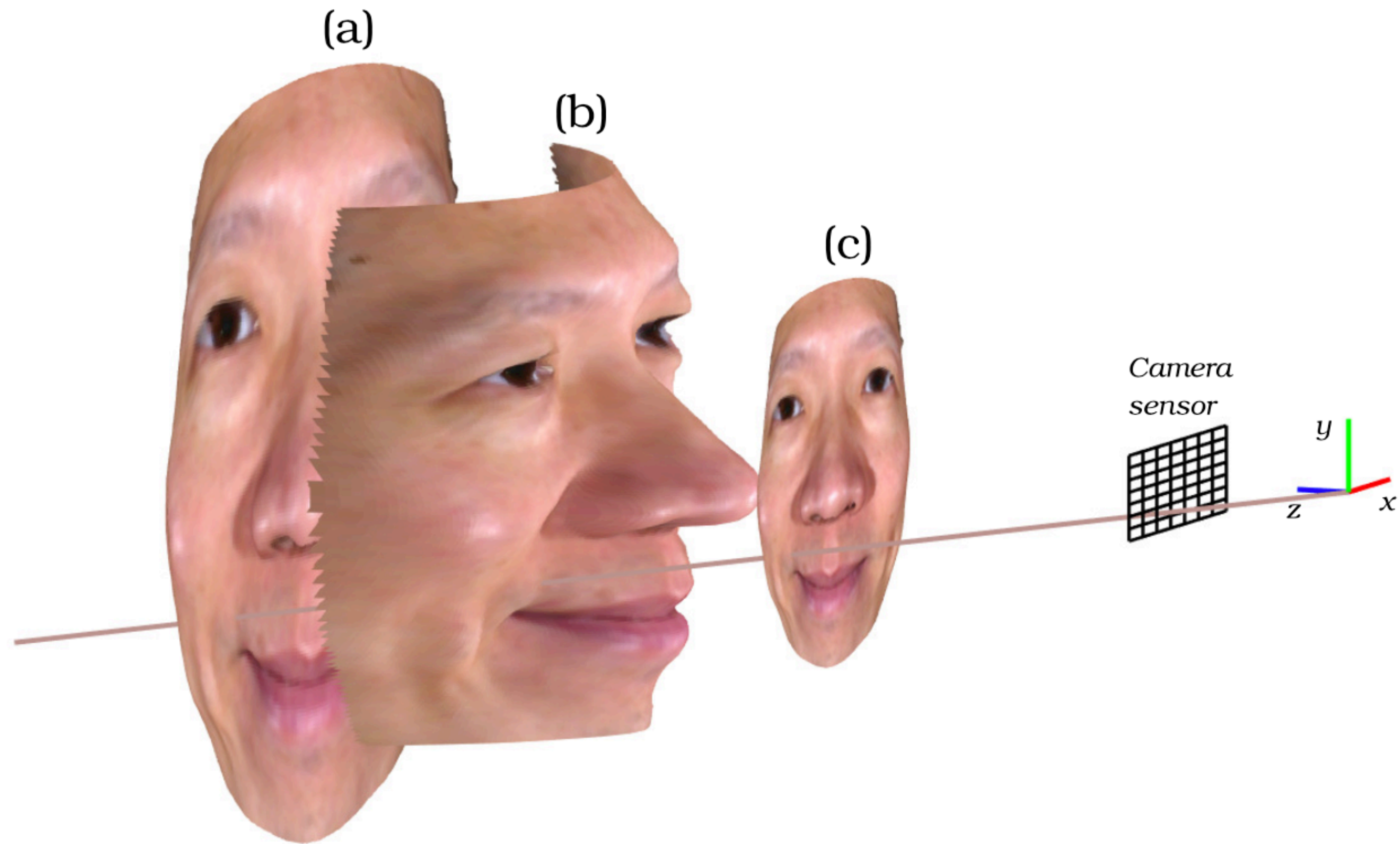


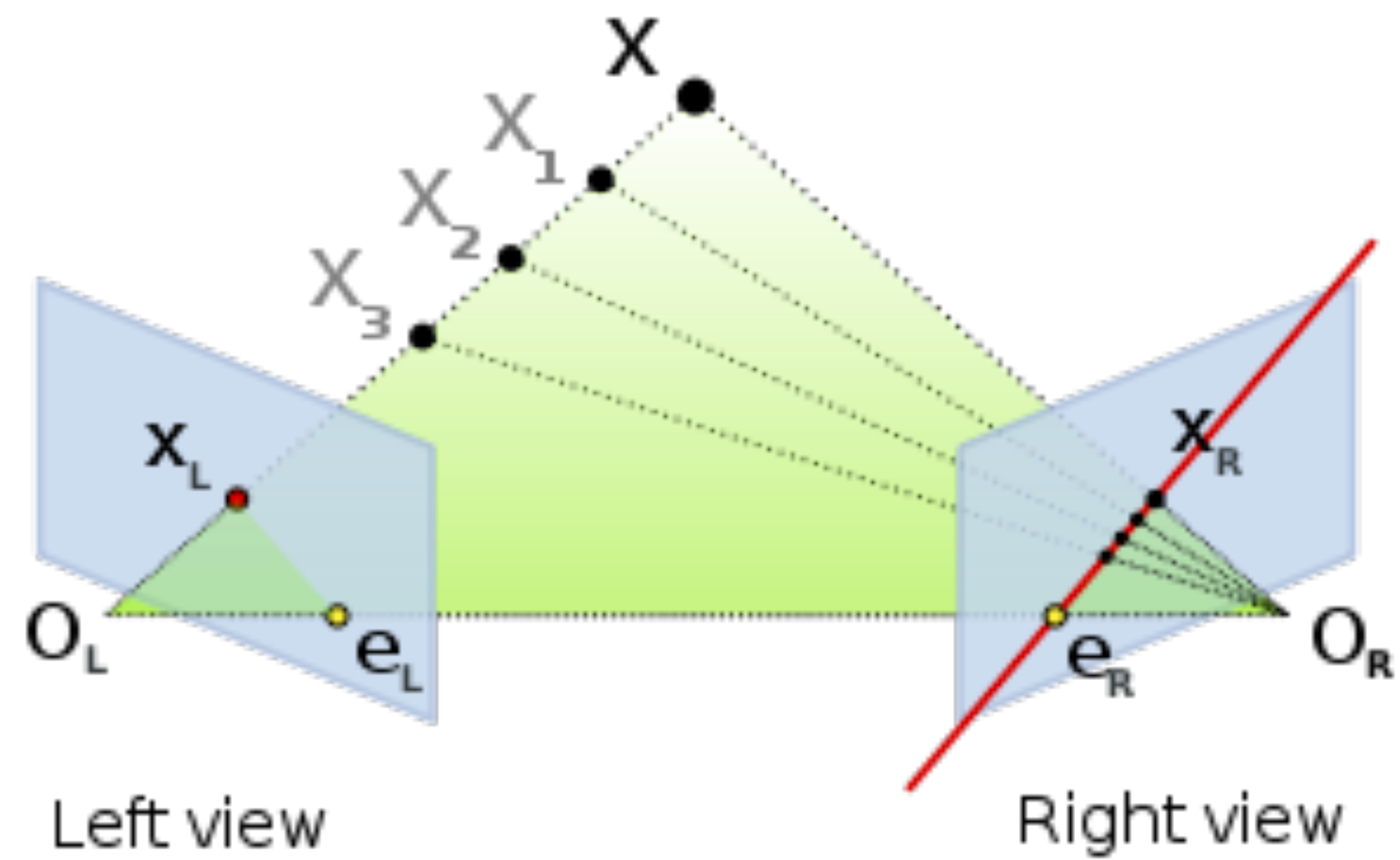
Image of surface (a)



Image of surface (b)

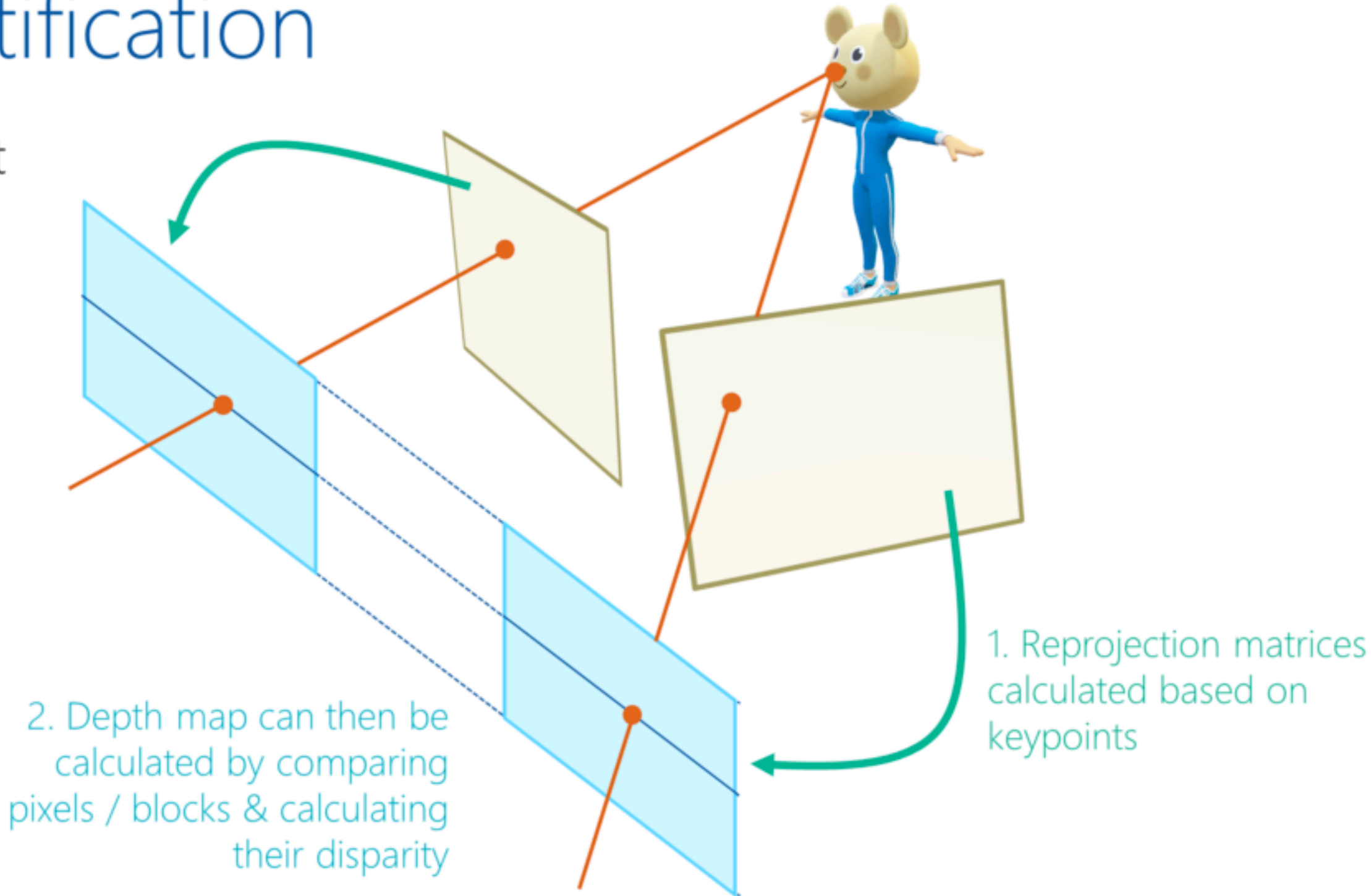


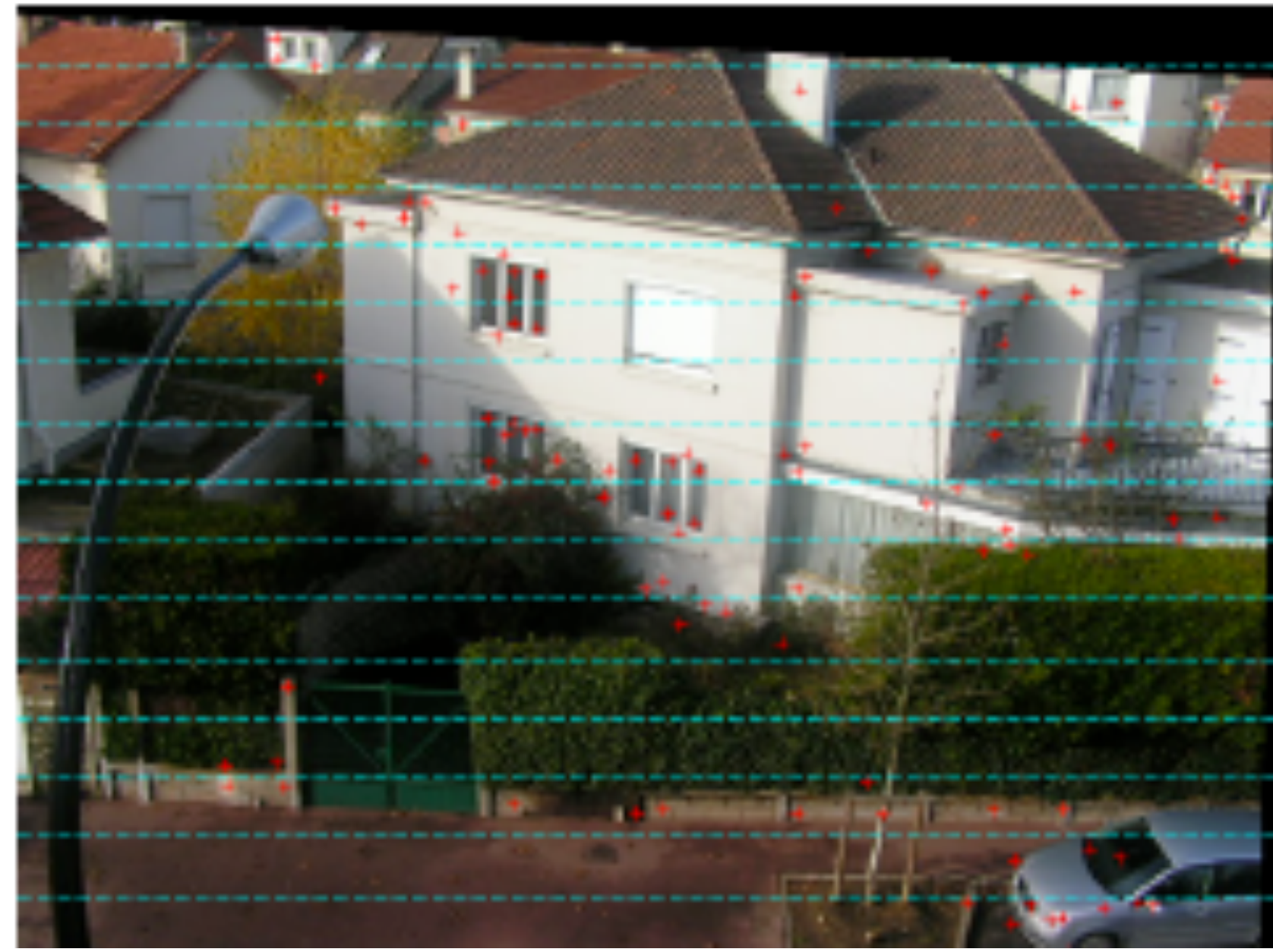
Image of surface (c)



# Stereo Rectification

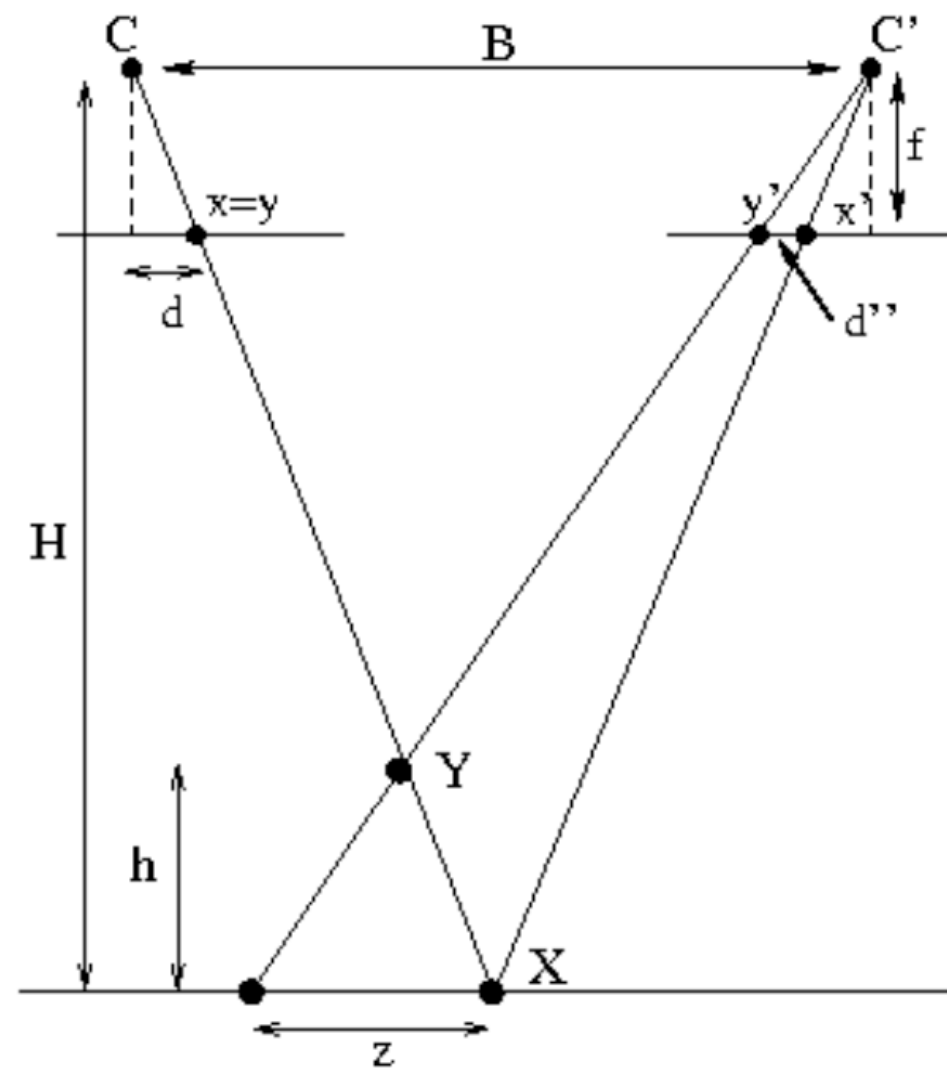
Reproject left & right image planes onto a common plane parallel to the line between camera centers





# Triangulation

## Fundamental principle of stereo vision



$$h \simeq \frac{z}{B/H}, \quad z = d'' \frac{H}{f}.$$

$f$  focal length.

$H$  distance optical center-ground.

$B$  distance between optical centers (baseline).

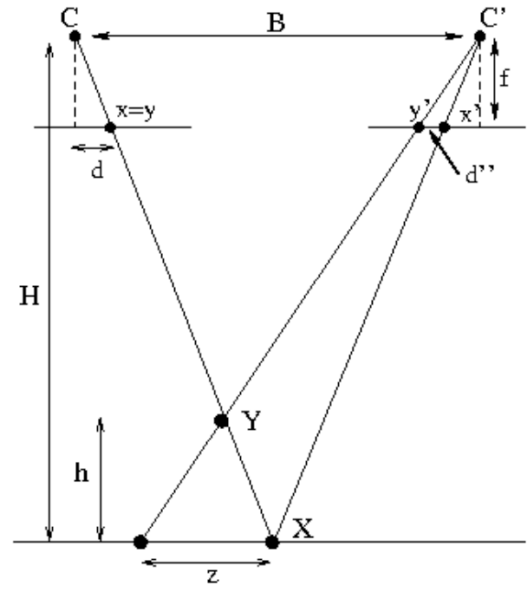
## Goal

Given two rectified images, point correspondences and computation of their apparent shift (disparity) gives information about relative depth of the scene.



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## Goal

Given two rectified images, point correspondences and computation of their apparent shift (disparity) gives information about relative depth of the scene.

- ▶ Let us write again the binocular formulae:

$$\lambda X = K(RX + T) \quad \lambda' x' = K' X$$

- ▶ Write  $Y^T = (X^T \quad \lambda \quad \lambda')$ :

$$\begin{pmatrix} KR & -x & 0_3 \\ K' & 0_3 & -x' \end{pmatrix} Y = \begin{pmatrix} KT \\ 0_3 \end{pmatrix}$$

(6 equations  $\leftrightarrow$  5 unknowns + 1 epipolar constraint)

- ▶ We can then recover  $X$ .
- ▶ **Special case:**  $R = Id$ ,  $T = Be_1$
- ▶ We get:

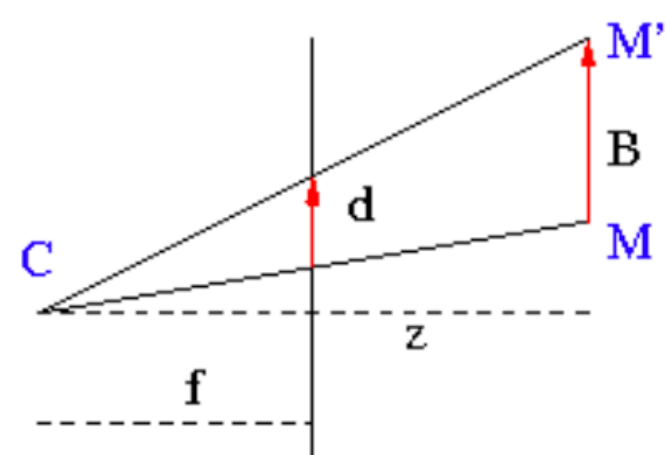
$$z(x - KK'^{-1}x') = (Bf \quad 0 \quad 0)^T$$

- ▶ If also  $K = K'$ ,

$$z = fB / [(x - x') \cdot e_1] = fB / d$$

- ▶  $d$  is the disparity

## Disparity map



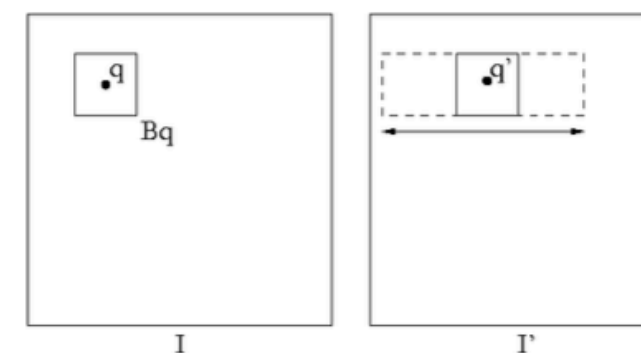
$$z = \frac{fB}{d}$$

Depth  $z$  is inversely proportional to disparity  $d$  (apparent motion, in pixels).

- ▶ **Disparity map:** At each pixel, its apparent motion between left and right images.
- ▶ We already know disparity at feature points, this gives an idea about min and max motion, which makes the search for matching points less ambiguous and faster.

## Local search

- ▶ At each pixel, we consider a context window and we look for the motion of this window.



- ▶ Distance between windows:

$$d(q) = \arg \min_d \sum_{p \in F} (I(q+p) - I'(q+de_1+p))^2$$

- ▶ Variants to be more robust to illumination changes:
  1. Translate intensities by the mean over the window.

$$I(q+p) \rightarrow I(q+p) - \sum_{r \in F} I(q+r) / \#F$$

2. Normalize by mean and variance over window.



Right image



Left image



Right image



Left image

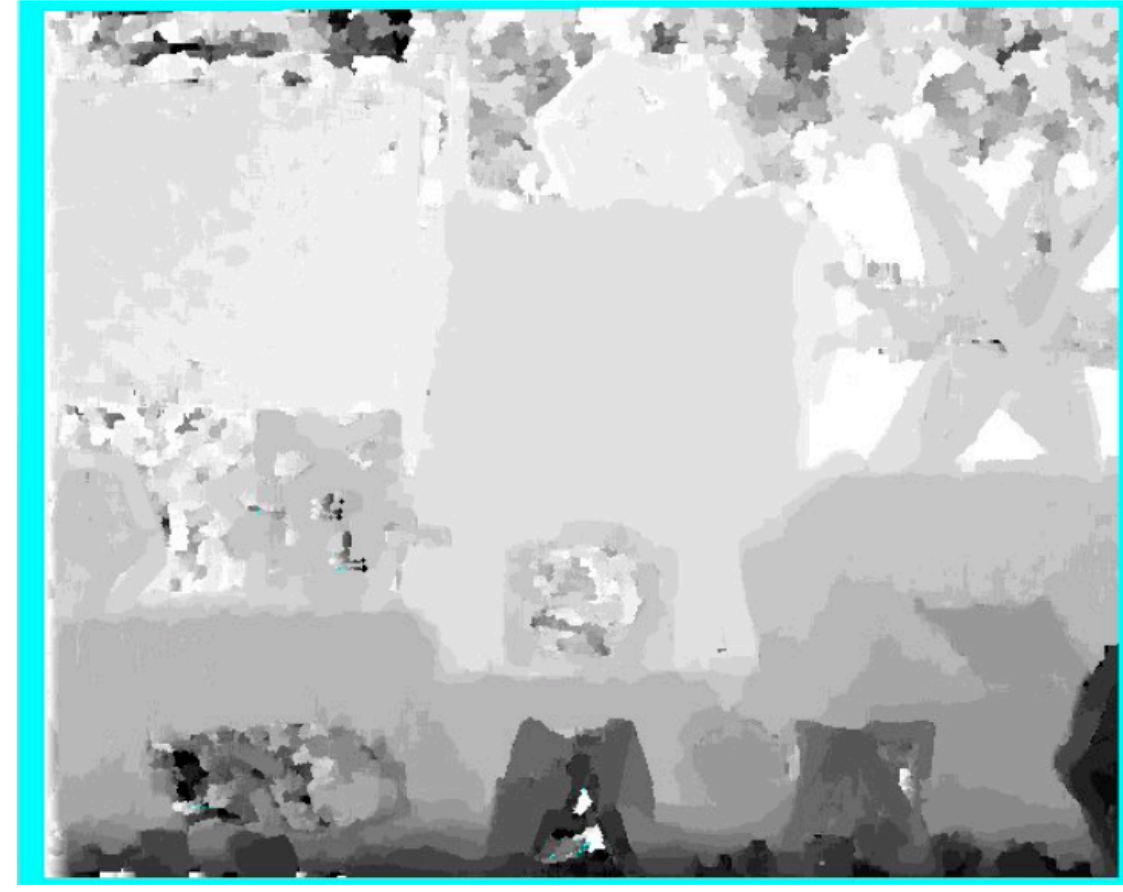




Right image



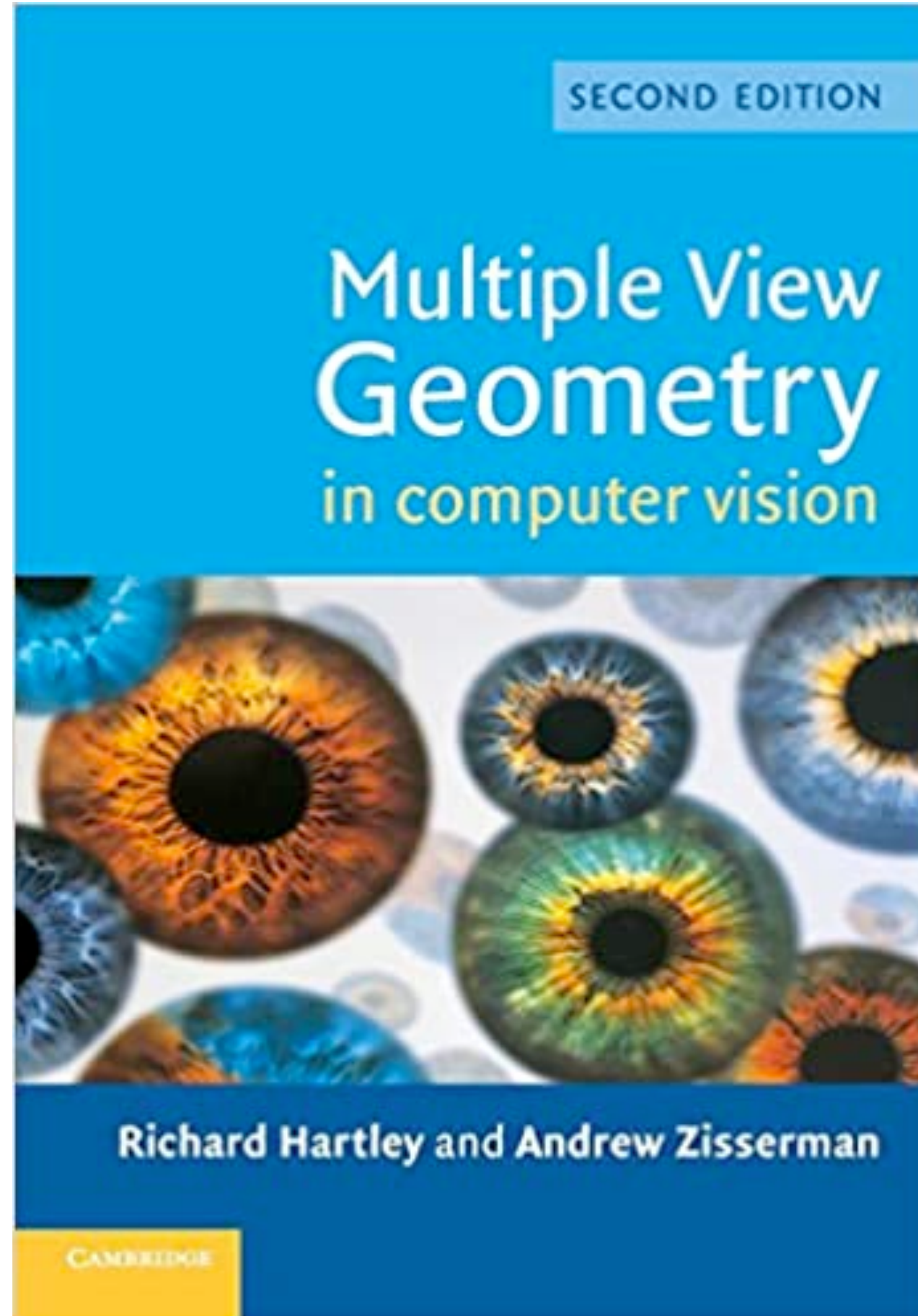
Left image



**Resumiendo**

- Modela una cámara y cuales son los parámetros más críticos.
- Pinhole model, matrix de proyección, parámetros intrínsecos y extrínsecos.
- Calibración de una cámara.
- Limites de vision monocular.
- Vision stereo puede usarse para obtener información 3D.
- Geometria epipolar.
- Disparidad entre imágenes epistolares contiene información de la profundidad.
- Desafíos asociados con la estimation de correspondencias (crucial para la calibración y para la estimación de correspondencias).

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## Quasi-Euclidean Epipolar Rectification

Pascal Monasse

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published · 2011-09-13

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*Communicated by* Luis Álvarez  
*Demo edited by* Agustín Salgado

### Abstract

The standard setup in reconstructing the three-dimensional geometry of a scene from a pair of stereo images is to have them rectified, in which case the apparent motion of points is horizontal. With pinhole cameras, it is always possible to find two homographies that rectify the images. The method of Fusiello and Irsara assumes that both cameras are the same with principal point at the center, but keeps the focal length as an unknown. The virtual rotations of the two cameras are then evaluated to minimize the vertical motion of points.


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software, datasets and online demo at  
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Pascal Monasse

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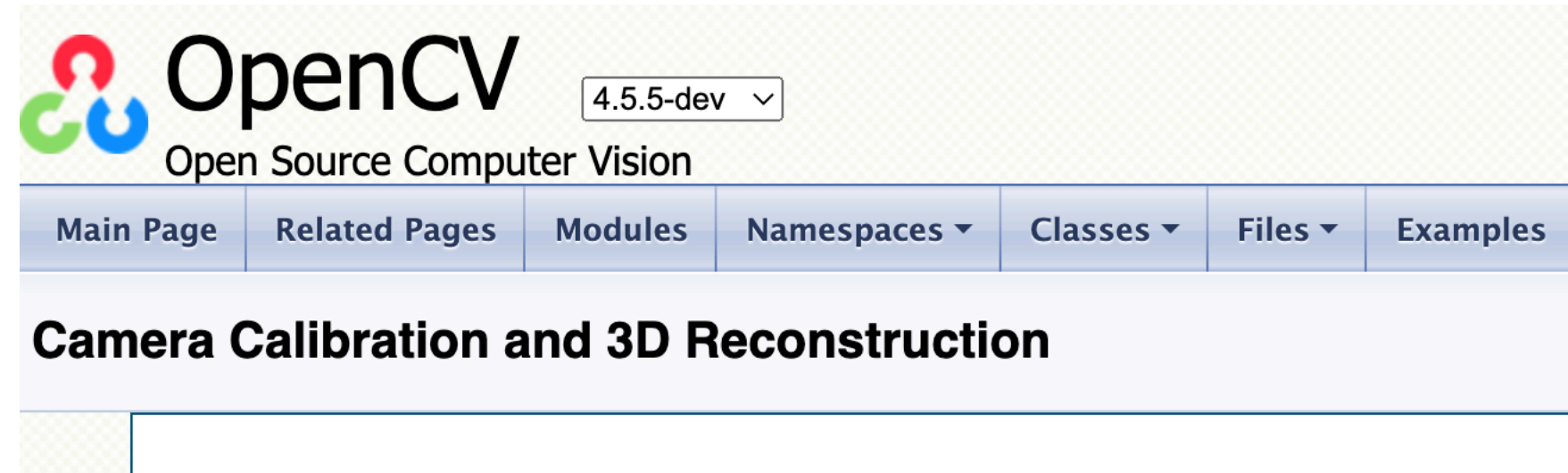
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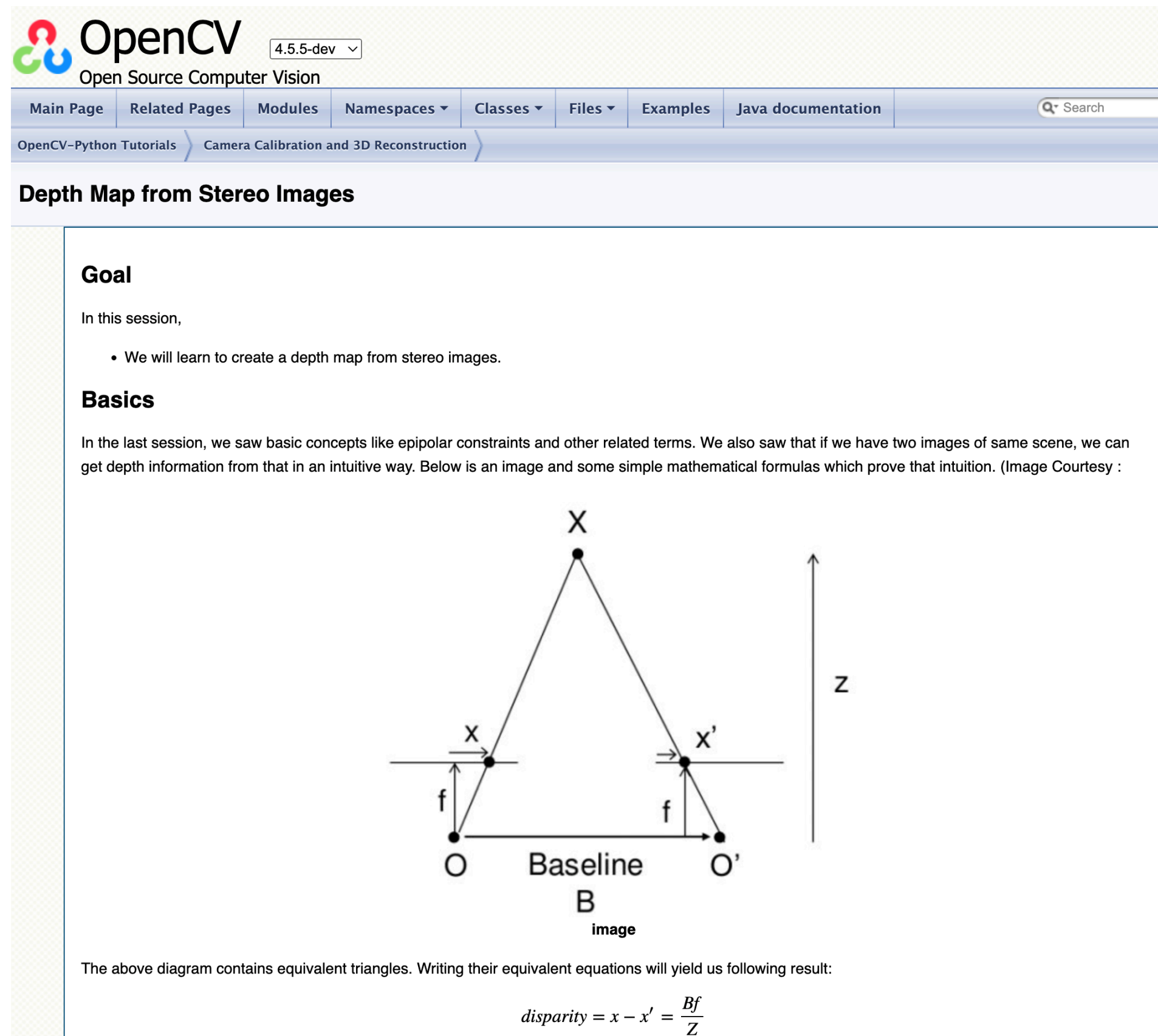
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# MATERIAL ADICIONAL



The screenshot shows the OpenCV website header with the logo and version '4.5.5-dev'. Below the logo is a navigation menu with links for 'Main Page', 'Related Pages', 'Modules', 'Namespaces', 'Classes', 'Files', and 'Examples'. The main heading is 'Camera Calibration and 3D Reconstruction'.



The screenshot shows the 'Depth Map from Stereo Images' tutorial page. It includes a 'Goal' section stating the objective is to create a depth map from stereo images. A 'Basics' section provides a recap of epipolar constraints and includes a diagram of a stereo camera setup. The diagram shows two camera centers, O and O', separated by a baseline B. A point X in the scene is projected onto the image planes at points x and x'. The focal length is denoted by f, and the depth of the point X is denoted by Z. The disparity is the horizontal distance between x and x'.

The above diagram contains equivalent triangles. Writing their equivalent equations will yield us following result:

$$disparity = x - x' = \frac{Bf}{Z}$$


Omar Padierna

Jan 2, 2019 · 10 min read · Listen



## Tutorial: Stereo 3D reconstruction with OpenCV using an iPhone camera. Part III.

Welcome to the third and final part of this [3 part tutorial](#) on stereo reconstruction.

A quick recap:

During the [first part](#) we covered a brief mention on the steps required for stereo 3D reconstruction and the gist of how stereo reconstruction works.



**Muchas gracias,  
Preguntas?**

**Gracias a Pascal Monasse por dejarme usar su material!**